

X-ray Imaging Micro-Calorimeter Spectrometer

Piet de Korte

On behalf of an emerging calorimeter collaboration



Netherlands Institute for Space Research

Contributors

GSFC(U.S)

Joe Adams, Simon Bandler, Regis Brekosky, Ari-David Brown,
Jay Chervenak, Megan Eckart, Richard Kelley, Caroline Kilbourne, Scott Porter, Jack Sadleir, Steve Smith

ISAS (Japan)

Kazu Mitsuda, Yoh Takei, Noriko Yamasaki

NIST(U.S)

Randy Doriese, Gene Hilton, Kent Irwin, Carl Reintsema, Joel Ullom, Leila Vale, *and others!*

PTB(Germany)

Joern Beyer, Dietmar Drung

SRON (Netherlands)

Marcel Bruijn, Bob Dirks, Luciano Gottardi, Henk Hoevers, Jan van der Kuur, Manuela Popescu, Marcel Ridder

VTT (Finland)

Mikko Kiviranta

SRON

Mirror Driven Specifications

- **Angular Resolution**

5 arc sec resolution = 485 - 606 µm for 20 – 25 m focal length

Proposed Pixel size between 250 – 300 µm

- **Field of View**

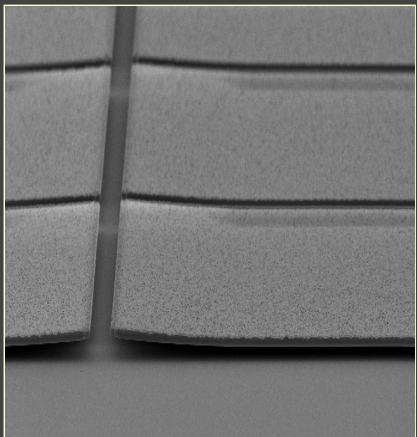
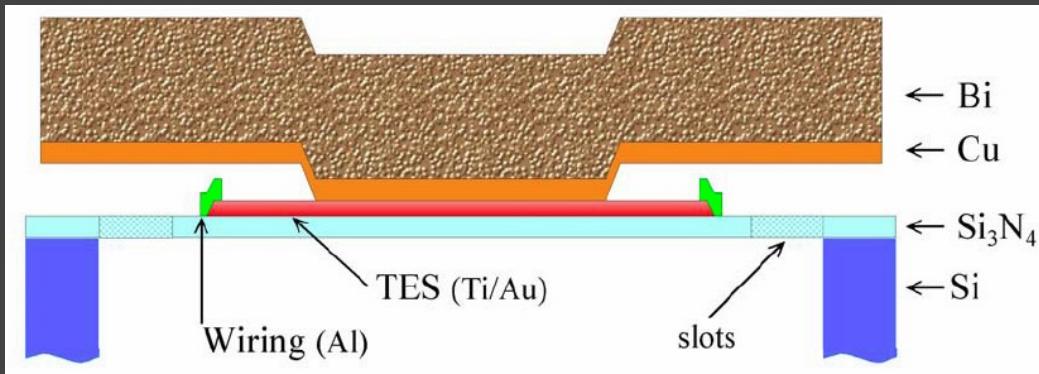
7 arc min radius = 71 mm

- **Countrate**

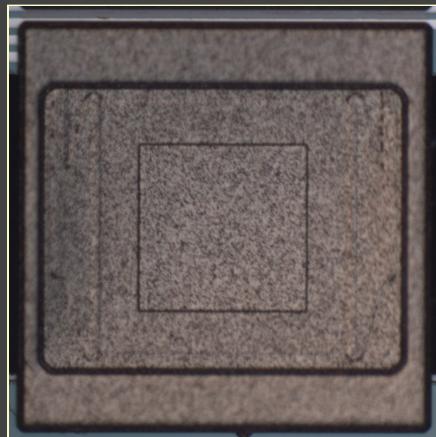
1mCrab ~ 125 c/sec (May 2008, NASA IXO mirror concept with $f = 20$ m)

TES-based Micro-Calorimeter

SRON PIXEL DESIGN

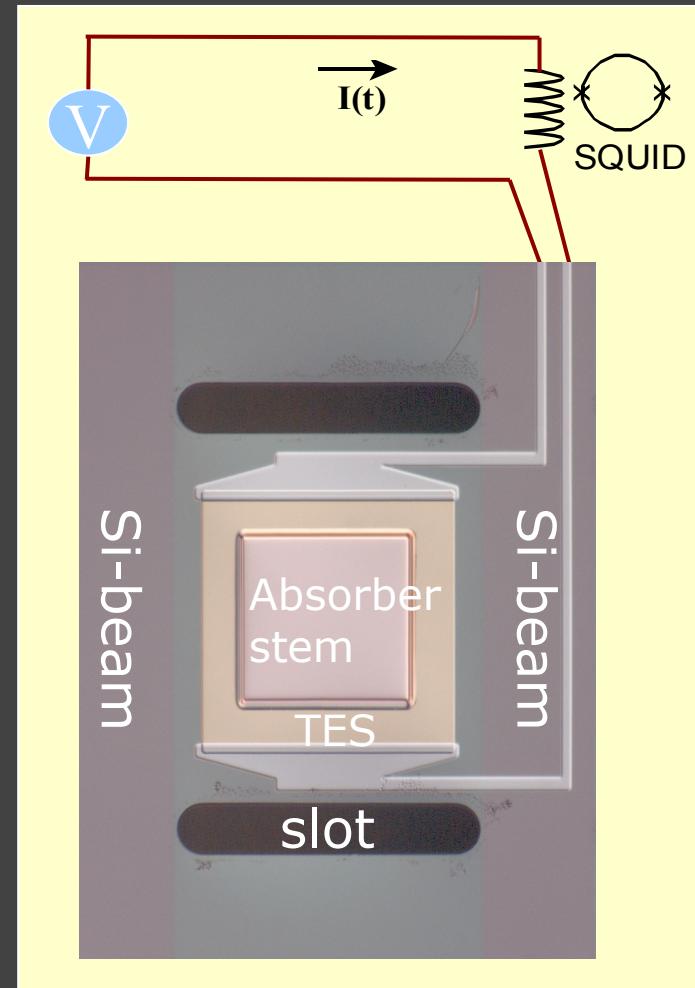


Side view



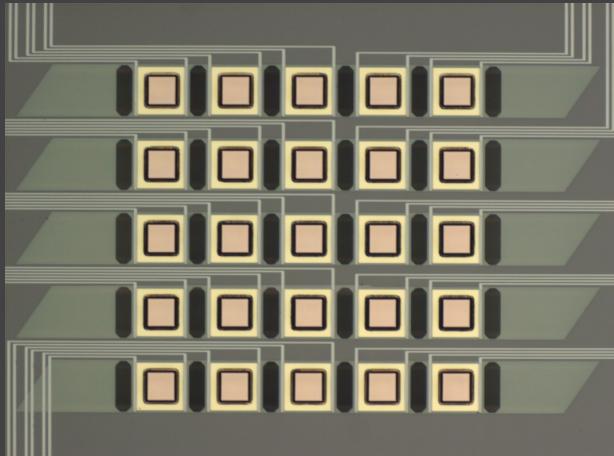
Top view

part of 5 x 5 array

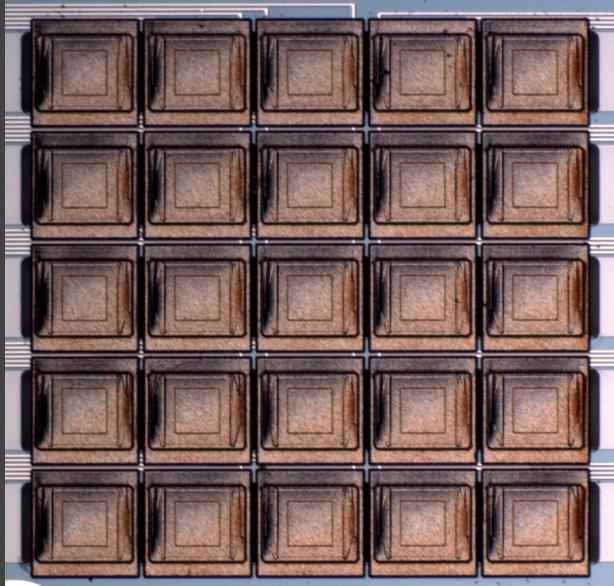


TES-based Micro-Calorimeter

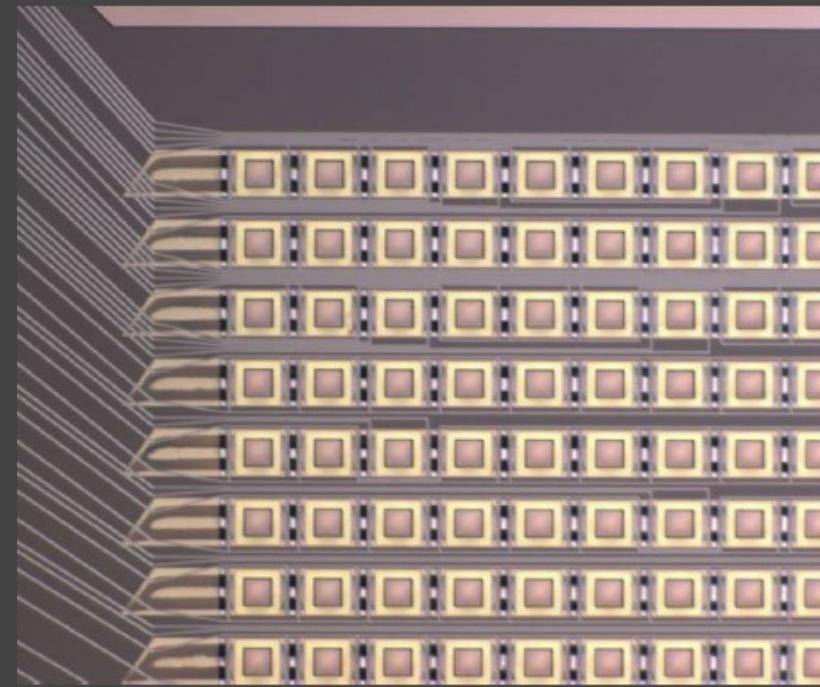
SRON ARRAYS



5 x 5 array
with Cu stems



5 x 5 array
with Cu/Bi
absorbers

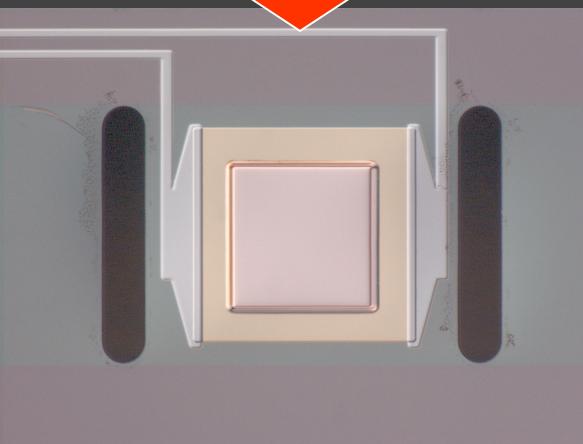


Close-up of 32 x 32 array

TES-based Micro-Calorimeter

PERFORMANCE for SRON PIXELS from 5×5 arrays

$$\Delta E_{TDL} \approx 3.1 \text{ eV } T_c = 105 \text{ mK}$$



Cu-absorber

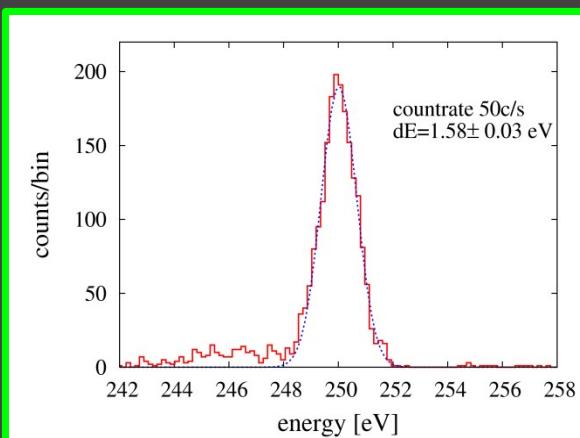
100 μs fall time

$\Delta E_{TDL} \approx 3.6 \text{ eV}$

$T_c = 116 \text{ mK}$

$$\Delta E_{TDL} = 2.35 \sqrt{k_B T^2 C}$$

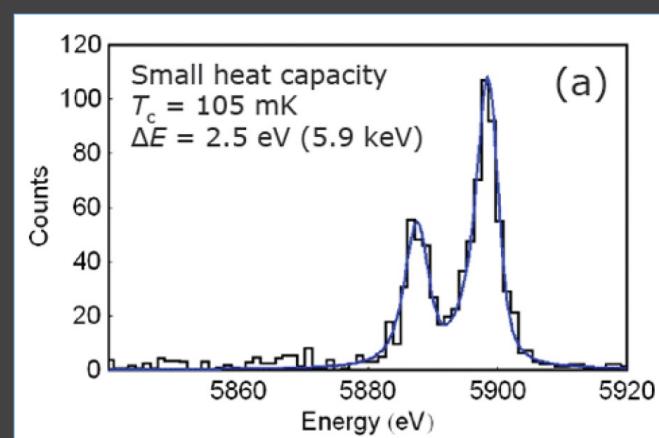
SRON



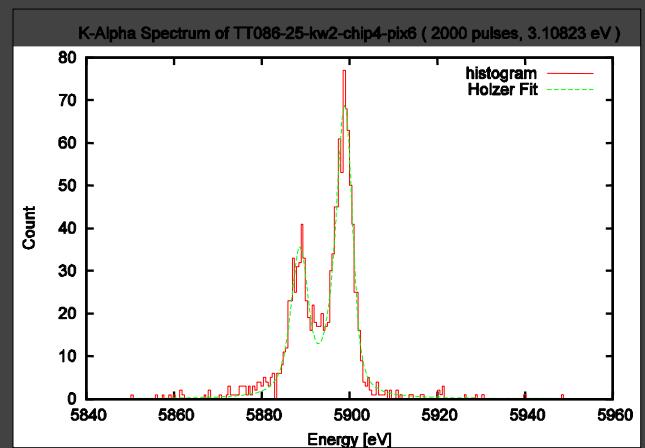
$\Delta E = 1.6 \text{ eV} @ 250 \text{ eV}$



Cu/Bi-absorber 0.3/3 μm



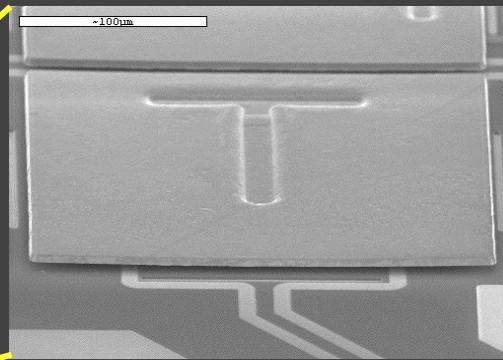
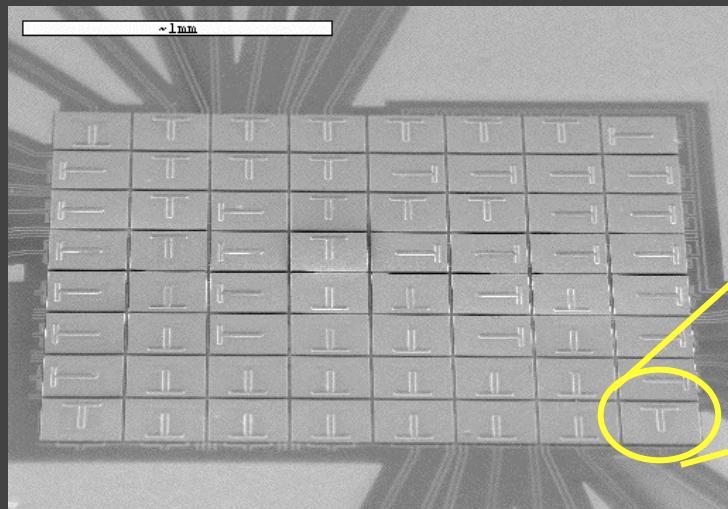
$\Delta E = 2.5 \text{ eV at 5.9 keV}$



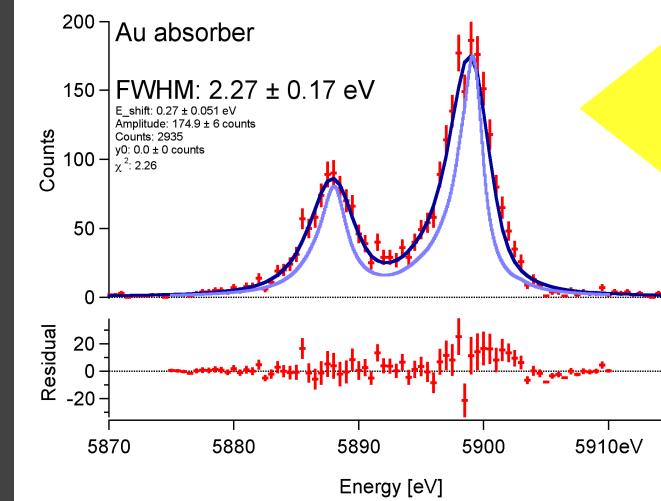
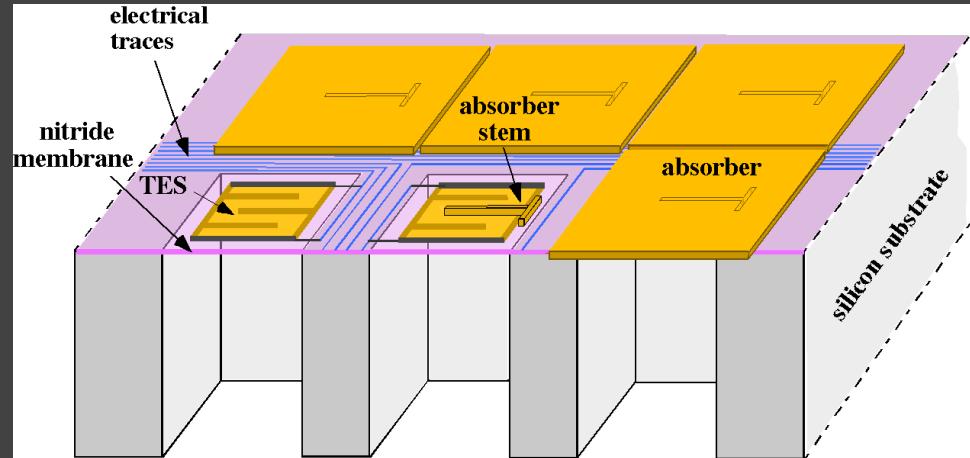
$\Delta E = 3.1 \text{ eV} @ 5.9 \text{ keV}$

GSFC TES approach

GSFC



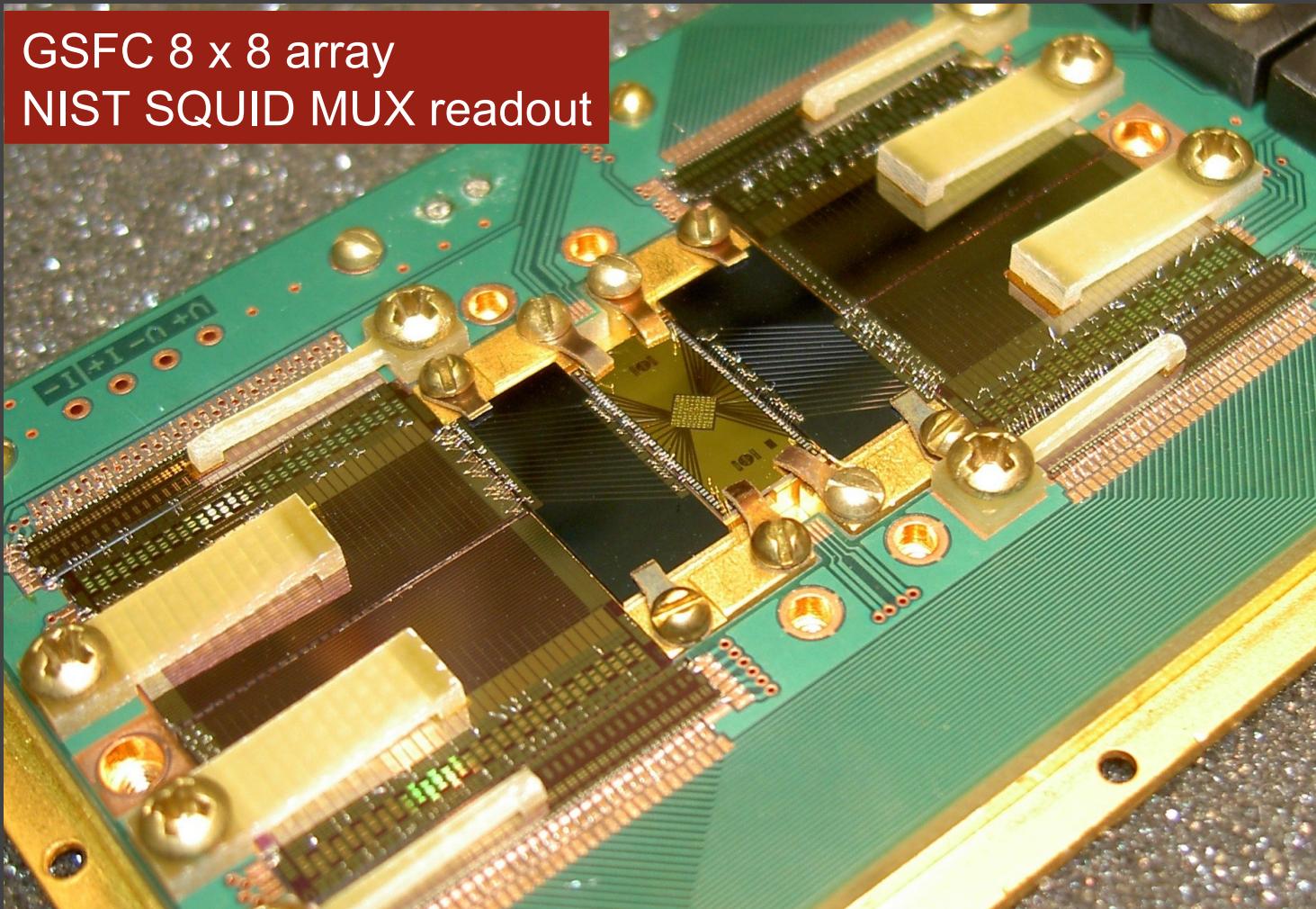
250 μ m



Multiplexed TES calorimeter array

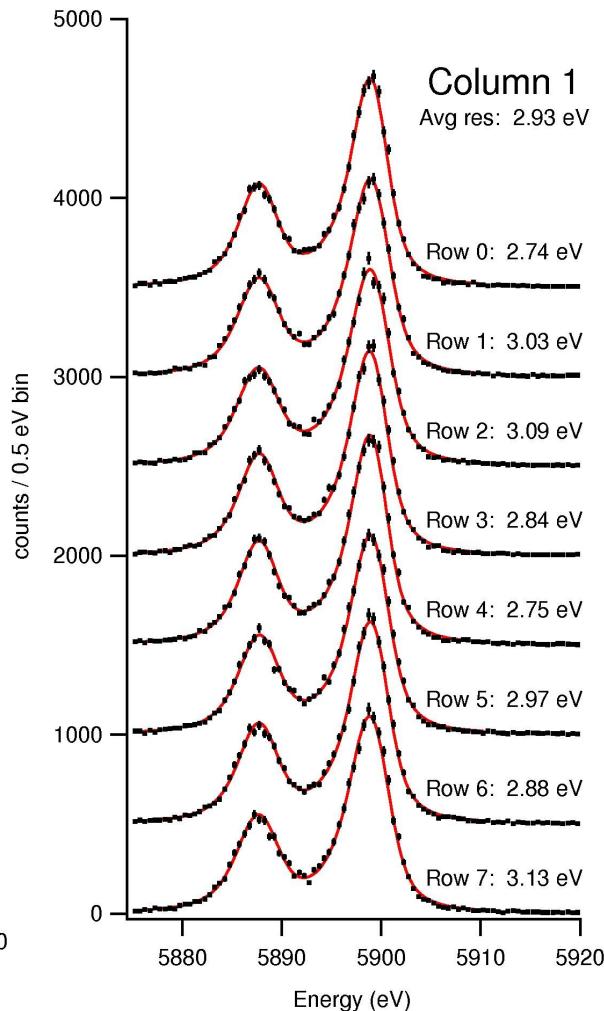
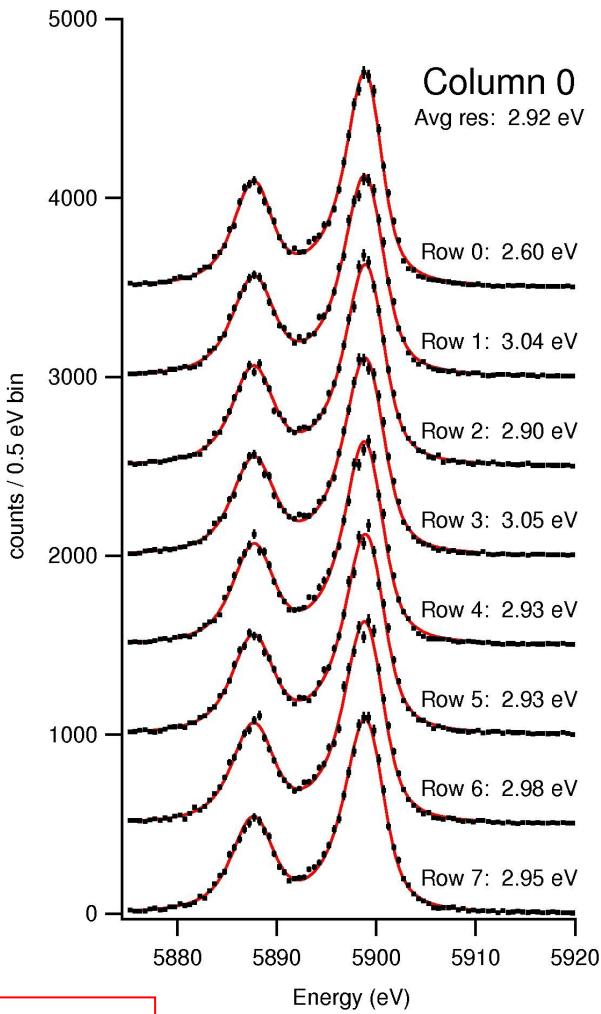
GSFC 8 x 8 array

NIST SQUID MUX readout



Also developed de-MUX software and we are now working on implementing real-time pulse height analysis

2×8 pixels read out with SQUID MUX



~30,000 counts per pixel from ^{55}Fe source
~500,000 total

$\tau_{\pm} = 280 \mu\text{ s}$
(critically damped)

2x8 MUX:
 $\langle \Delta E_{\text{FWHM}} \rangle =$
 $2.93 \pm 0.02 \text{ eV}$

GSFC

SRON

August 20-22, 2008

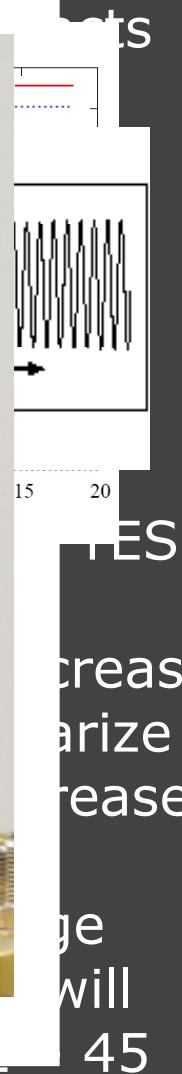
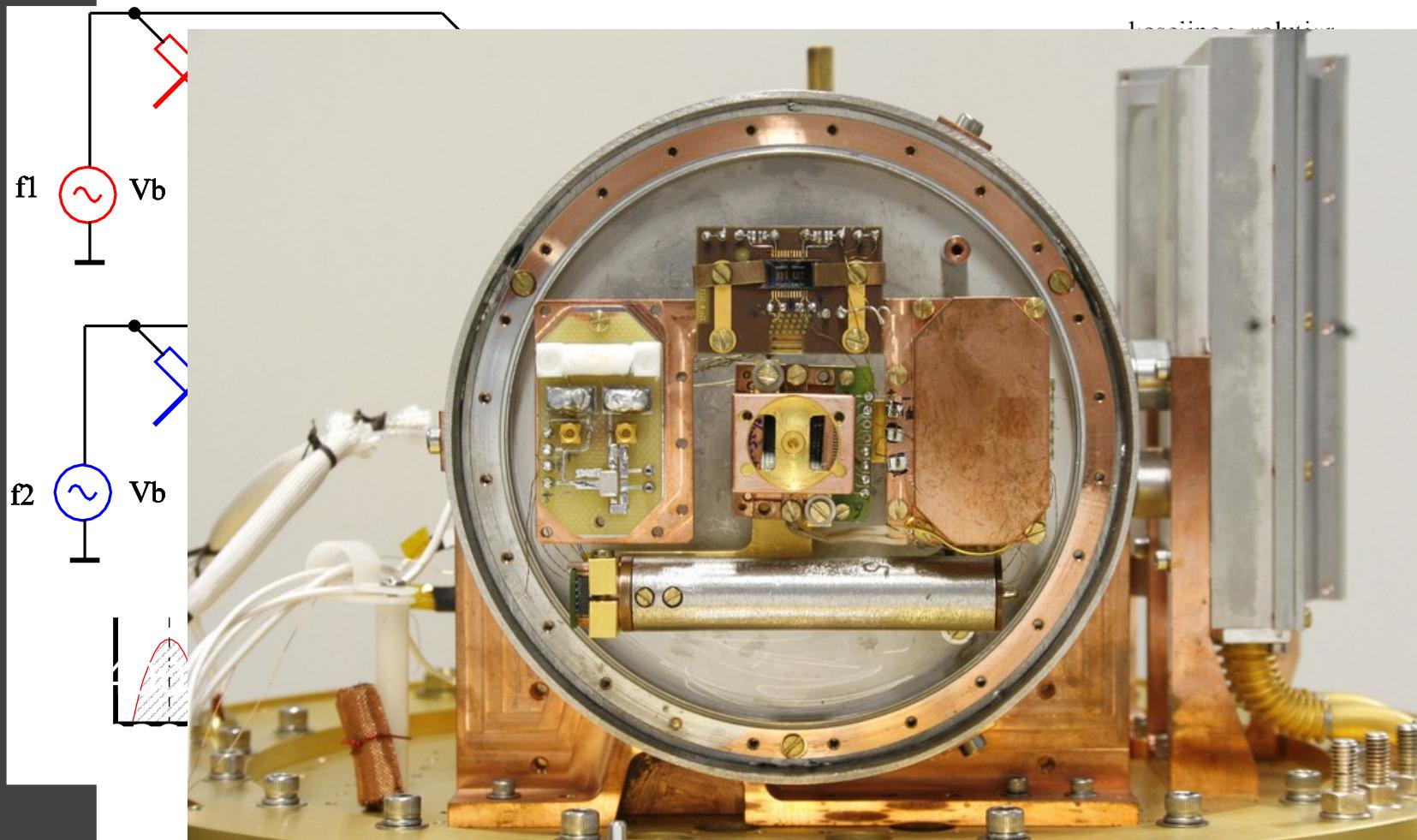


National Institute of Standards and Technology

Facility Science Team - GSFC

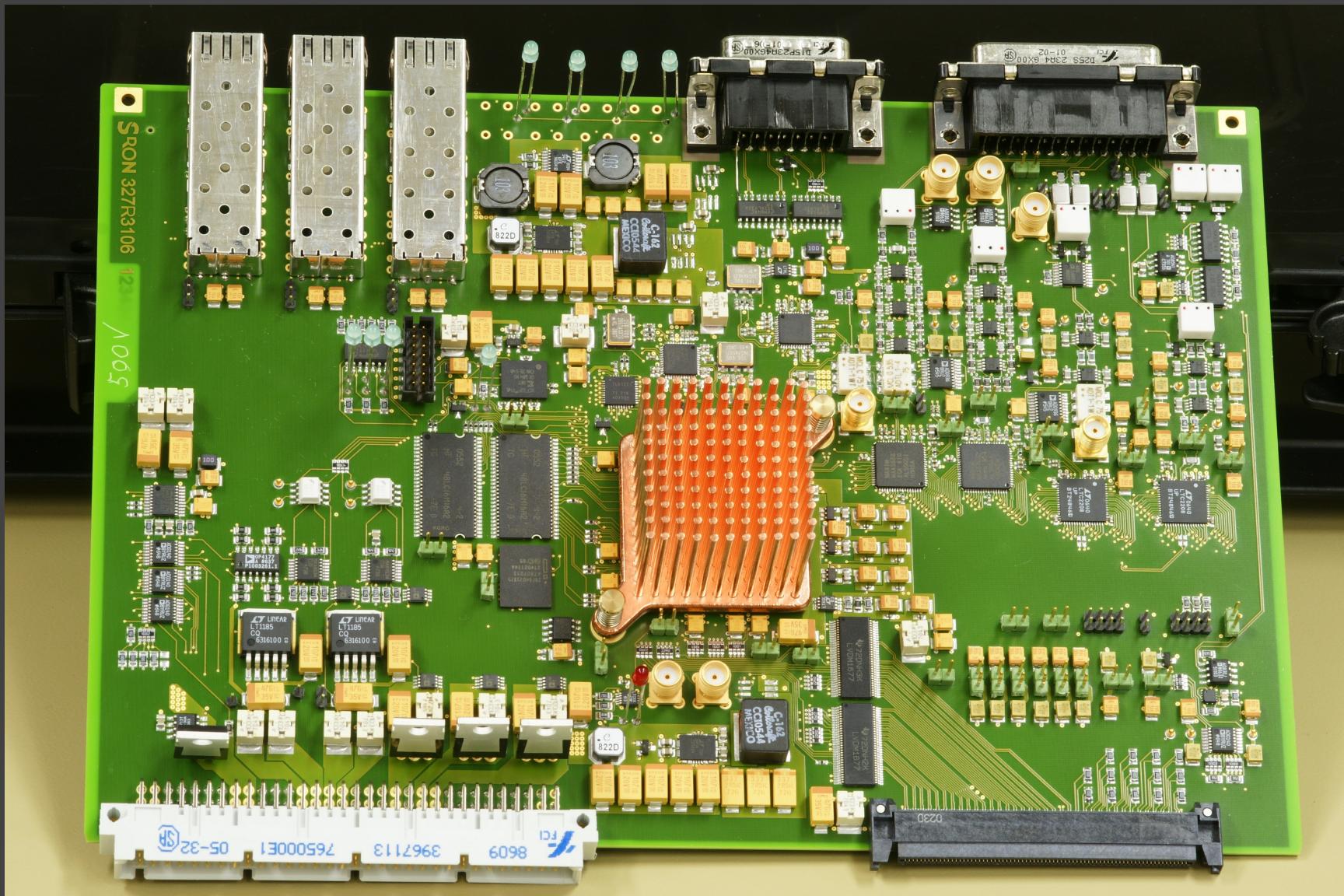
IXO Meeting Munich 17 - 19 September 2008

FREQUENCY DOMAIN MULTIPLEXING CURRENT SUMMING TOPOLOGY

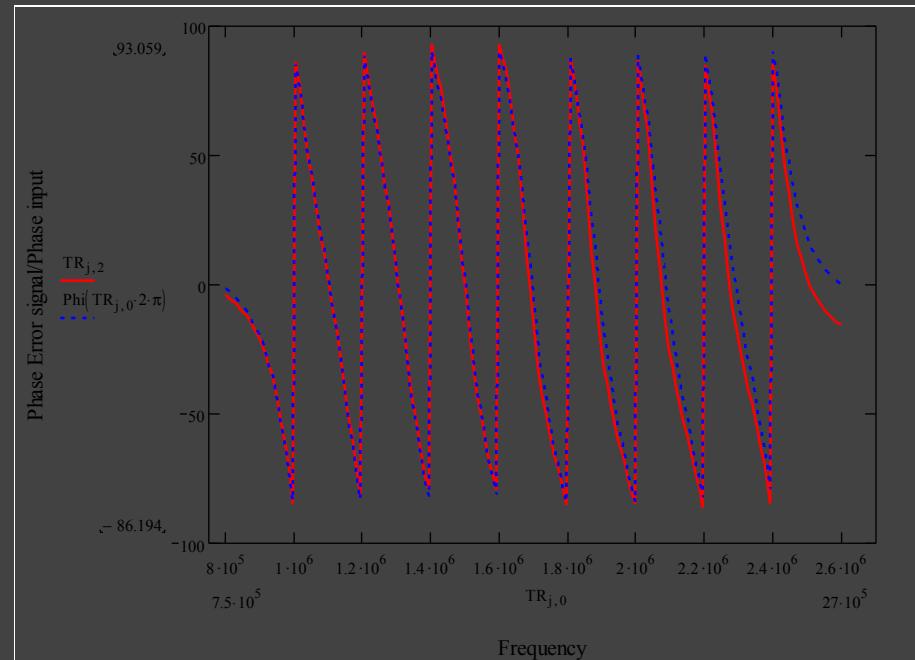
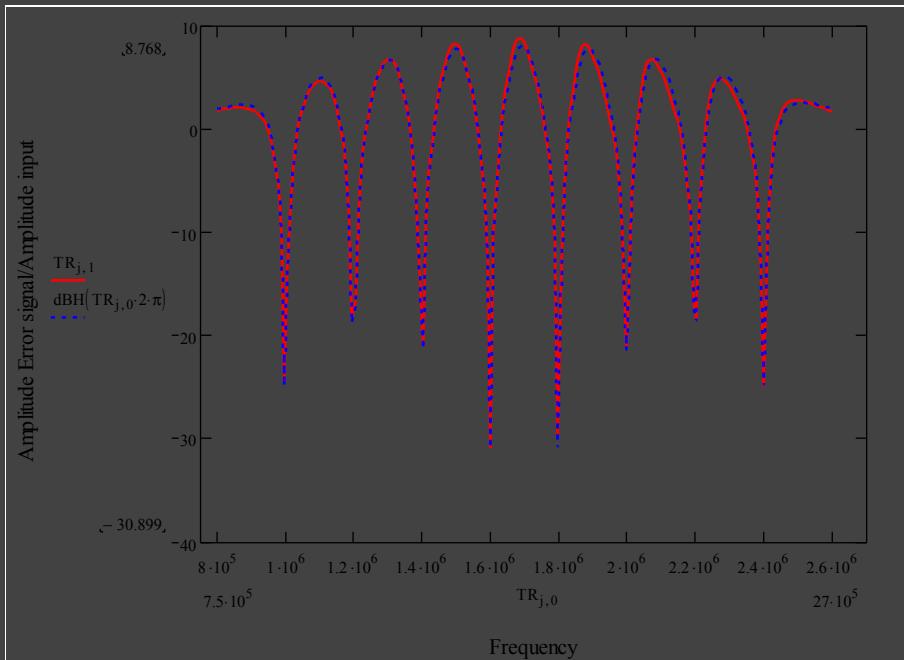


enable multiplexing of 32 pixels/channel

BBFB electronics board realization



Amplitude and Phase measurements/model of BBFB On a commercial Xilinx breadboard



Amplitude: red-data blue-model

Phase: red-data blue-model

Gain-bandwidth of 35 kHz for 200 kHz spacing and 830 ns delay

FLL-gain of 3.5x at highest signal frequency (10 kHz) and 22 x at 1.6 kHz (100 μ s pulse decay time)

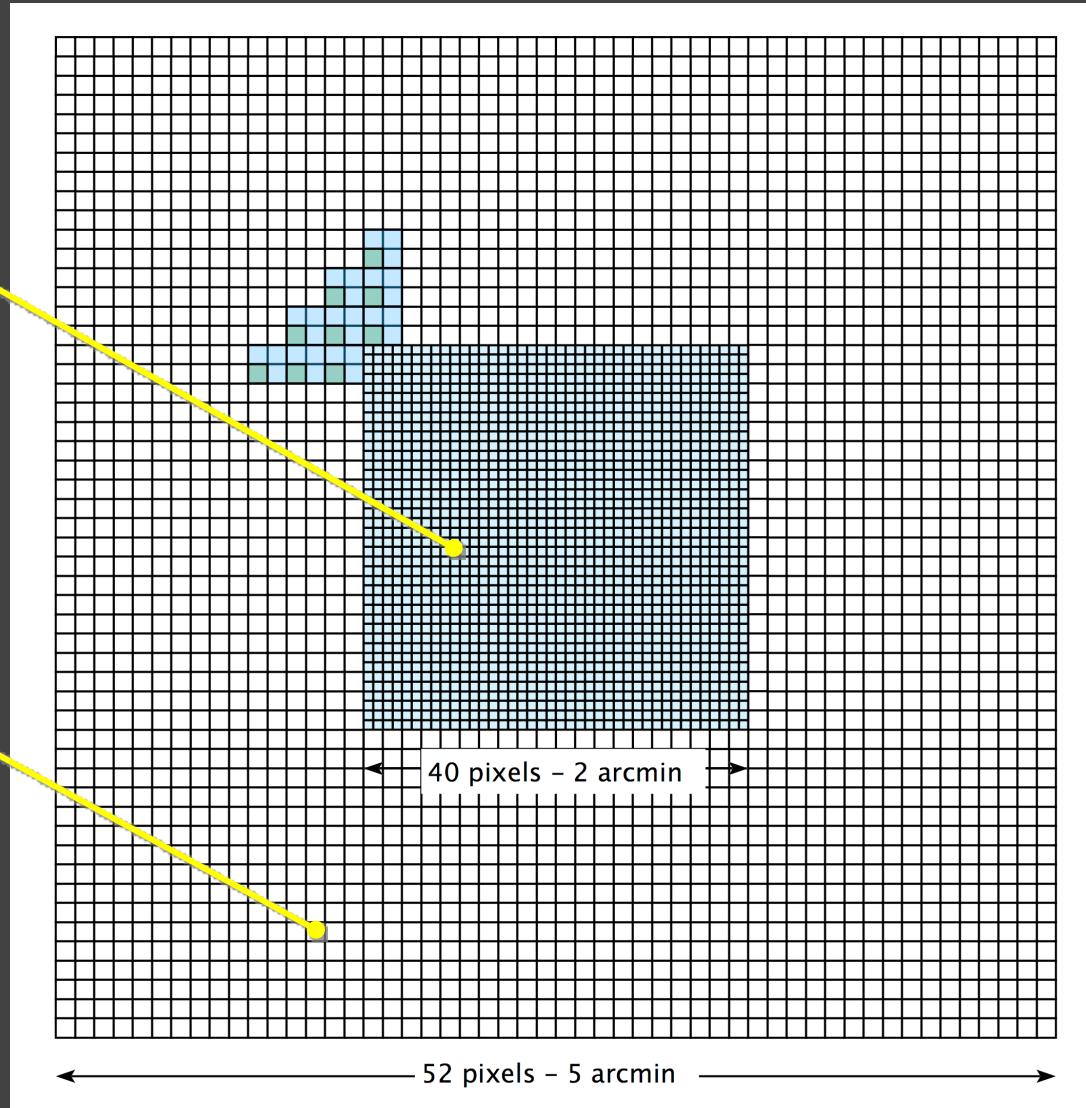
Focal Plane Array Layout (from Con-X→IXO)

Central, core array:

- Individual TES - one absorber/TES (40 x 40)
- 2 arcmin FOV
- 2.5 eV resolution (FWHM)
- Fast (< 300 μ sec time constant)

Outer, extended array

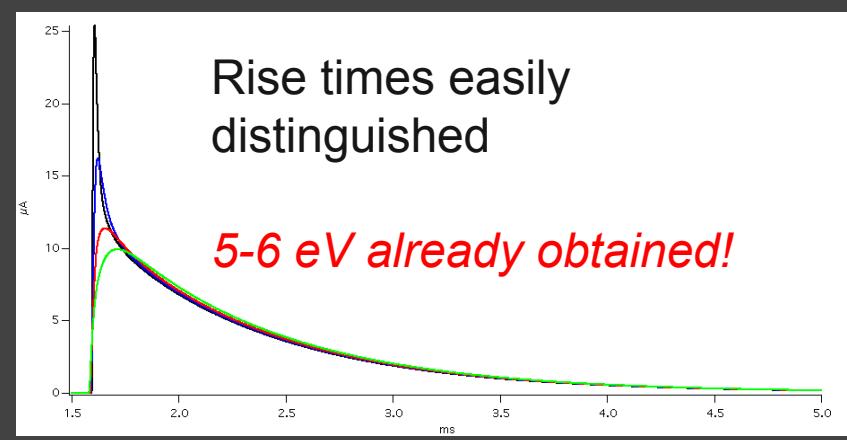
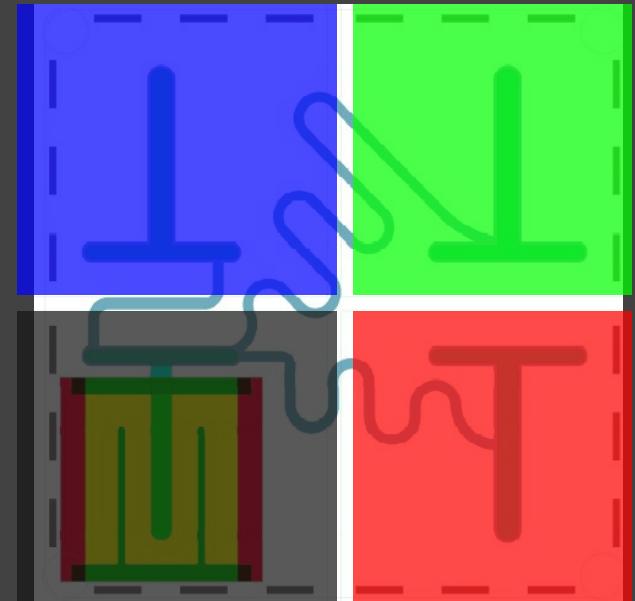
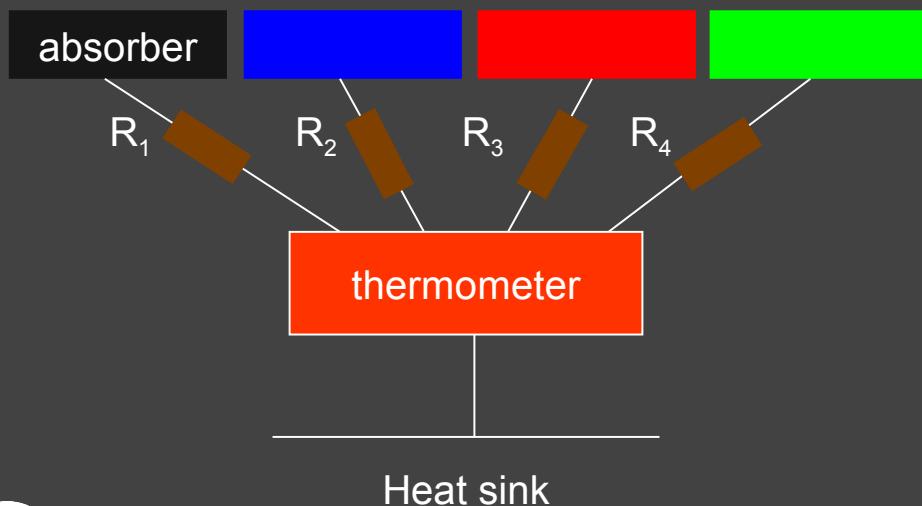
- 4 absorbers/TES
- Extends array to 52 x 52 pixels for a total of 2176 readout channels
- 5.0 arcmin FOV
- < 10 eV resolution
- ~ 2 msec time constant



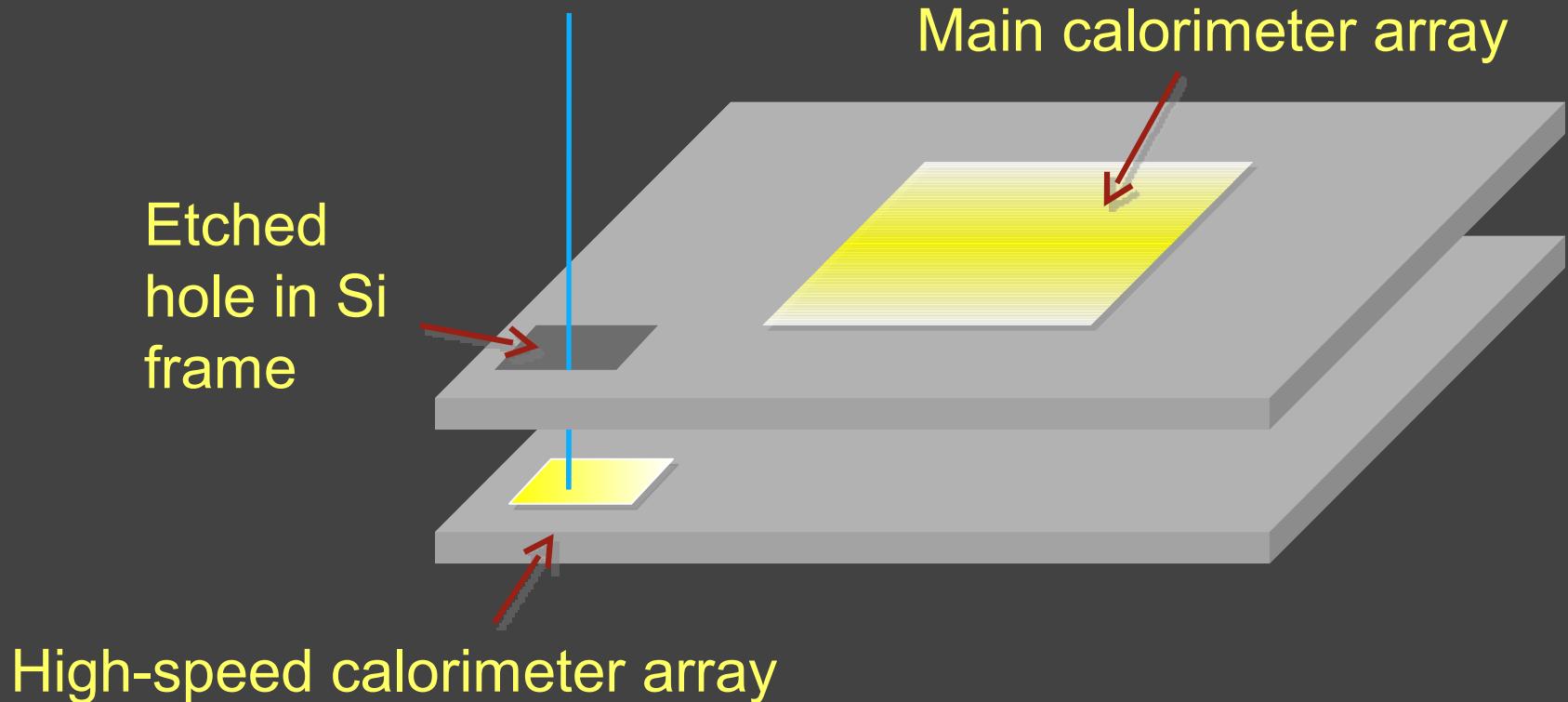
Multi Absorber TES - 1 TES, 4 absorbers

Simple approach:

Separate absorbers (e.g., 4) connected to a single TES, each with a different thermal conductance.



Optimized high-speed array (GSFC)



- 20 x 20 array of 1 arcsec pixels
- Distribute counts over ~ 10 times more pixels
- Use direct coupling to Si substrate for higher speed (~ 10's of micro-sec.)

ISAS/JAXA COOLING CHAIN

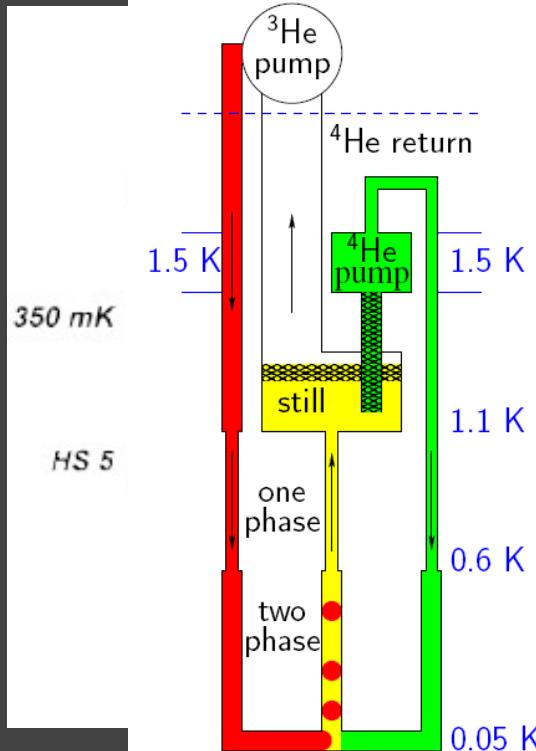
Cooler	1ST (100K)	2ST (20K)	2ST+ ⁴ He JT (4K)	2ST+ ³ He JT (2K)
				
Specification	2W@80K 50W, 4.2kg	325mW@20K 90W, 9.5kg	20mW @4.5K 120W, 23kg	16mW@1.7K 190W, 25kg
Ground test status	Life time test > 5 years (still running)	Life time test > 4 years (still running)	1 year test was done. A new lifetime test in preparation	Lifetime test in preparation
Mission status	Suzaku, in orbit 3.1 years	Akari, in orbit 2.5 years	FM for SMILES assembled	EM for SPICA & Astro-H(NeXT) assembled

Last stage cooler developments in Europe

Interface with satellite cryostat at 2.5 K with 10 mW cooling power

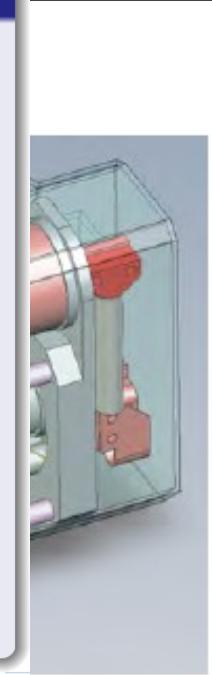
Options under development:

Closed He 3 option / last stage ADR (CEA Grenoble) XANeel
2 stage option / last stage ADR (Air Liquide, CNES, SODENA)
30 W and 31 kg for 1 μW during 30 hours
25 W and 5 kg for 1 μW during 30 hours



XEUS - NFI X-ray experiment

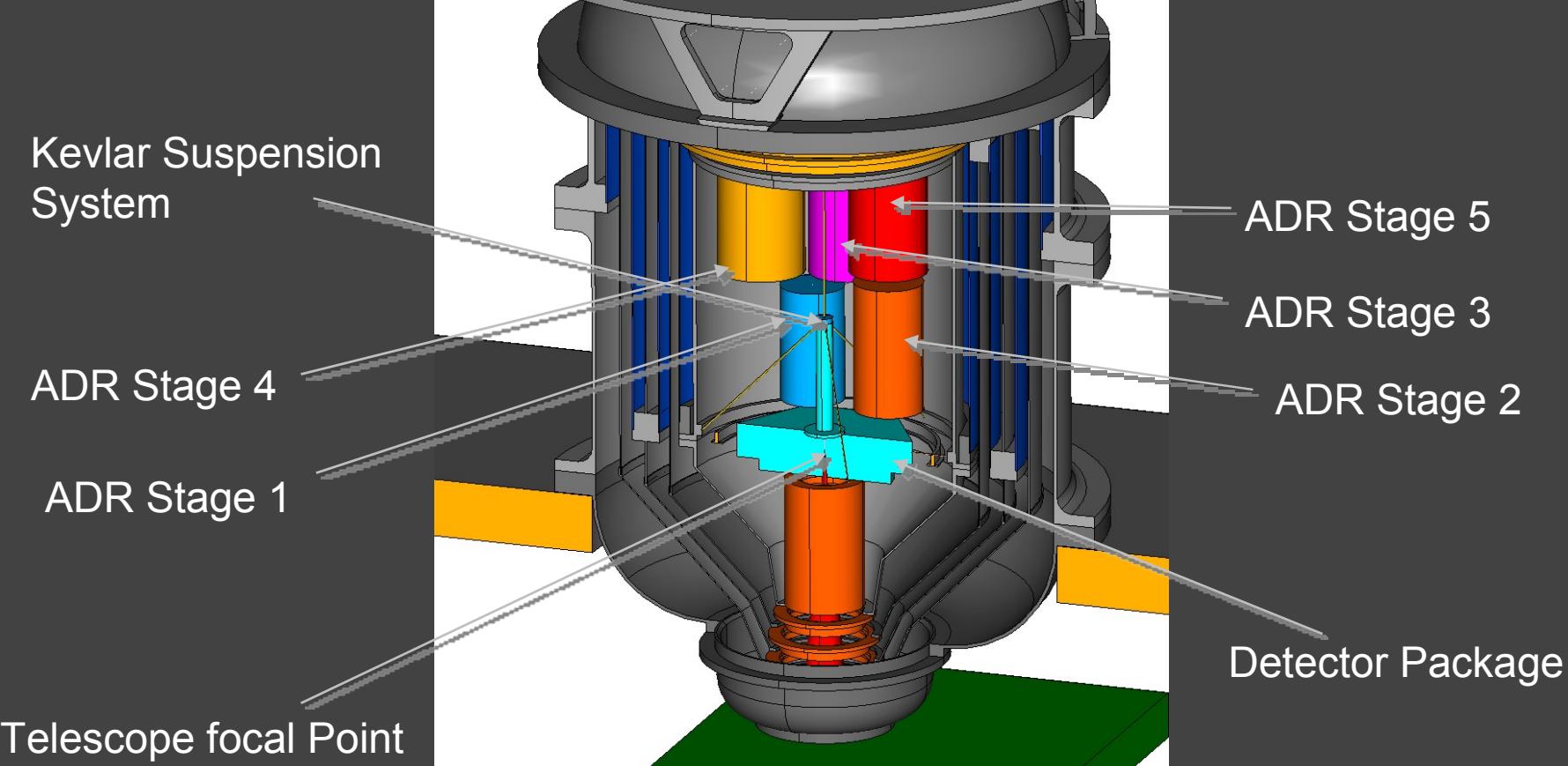
- requires 1 μW at 50 mK for ADR
 - CCDR meets requirement with $\dot{n}_3 = 30 \mu\text{mole/s}$, $\dot{n}_4 = 120 \mu\text{mole/s}$ and a heat exchanger of $L = 9 \text{ m}$ and $d = 0.4 \text{ mm}$
 - better thermalization of wiring reduces required cooling power for CCDR
- precooling stage of XEUS delivers 10 mW at 2.5 K
 - CCDR needs 5 mW at 1.5-1.8(?) K
 - solution: $^3\text{He Joule-Thompson}$ expansion from 15 K or with SPICA technology



AIR LIQUID

cnes

Cryostat design adopted for recent IDL study at GSFC



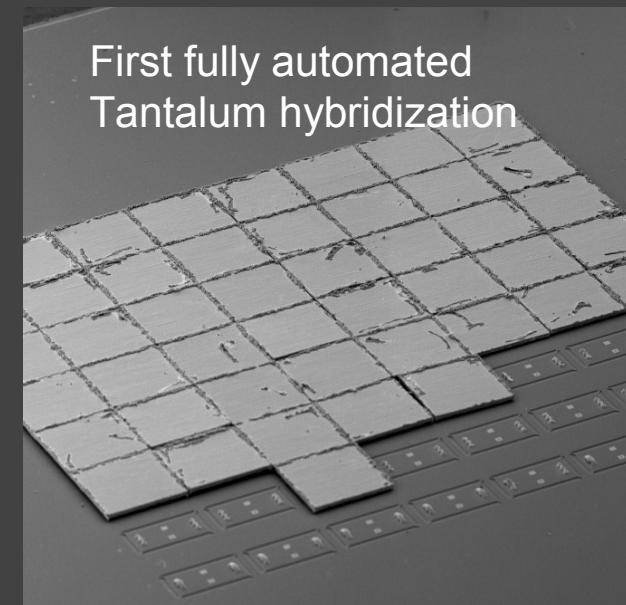
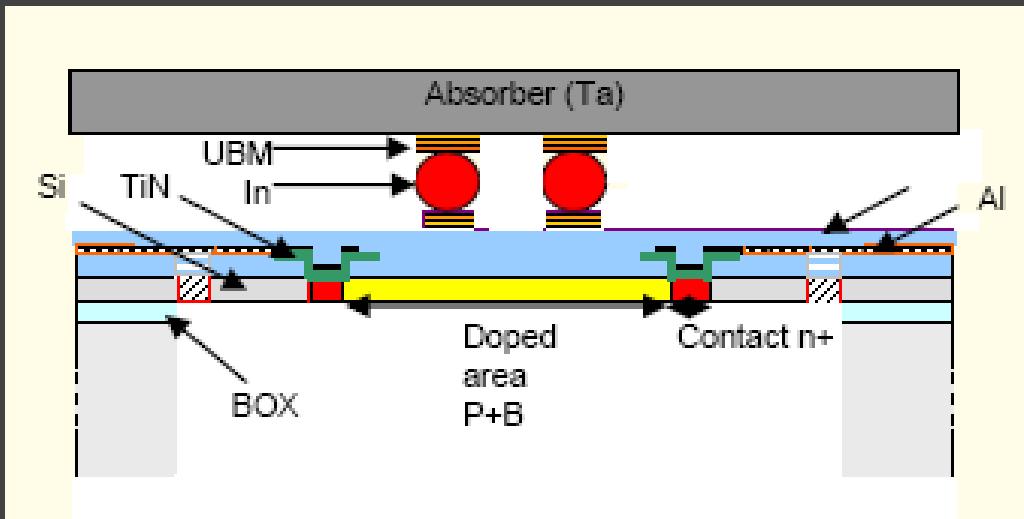
Facility Science Team - GSFC

Si-doped X-ray Micro-calorimeter at CEA-Saclay

Herschel heritage: Developments by CEA-Saclay and LETI, Grenoble

Contributed paper by Claude Pigot

Fully integrated sensor with read-out multiplexer



Results: - Impedance of 8X8 sensor matrix in the right range with good sensitivity

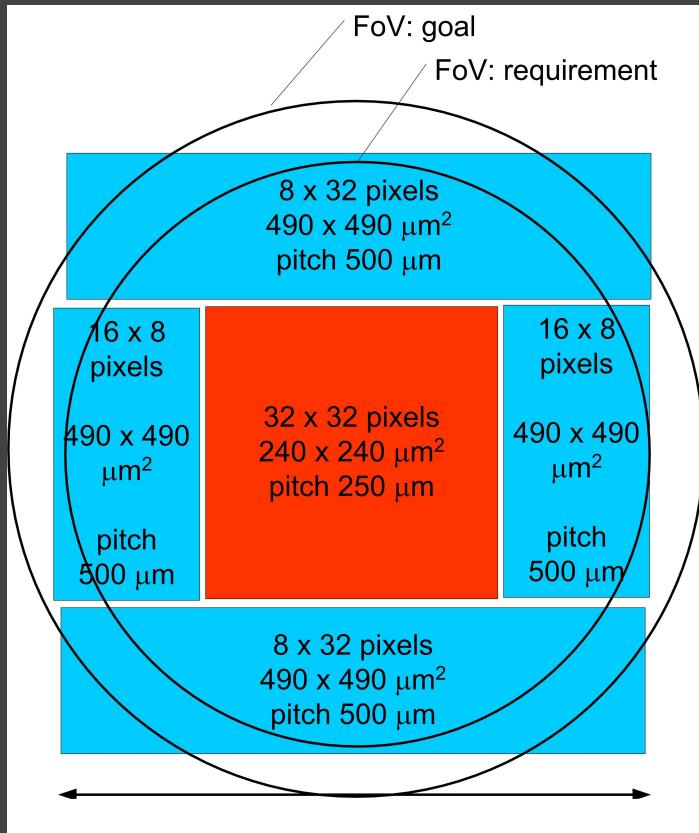
- Integration of absorber matrix onto sensor matrix promising

Next steps: April 2008: First 8X8 array with freed Sensor & Absorber
 End 2008: 1st Iteration Cold Electronics

Pro: Fully integrated system with multiplexed read-out

Con: Till now no X-ray performance data, use of Ta-absorbers by other teams failed, potentially slow response, developments late for XEUS.

Focal Plane Array Layout for XEUS → IXO



Field of View: $2.75 \times 2.75 \text{ arcmin}$

Central pixels: $1.37 \times 1.37 \text{ arcmin}$

$2.5 \times 2.5 \text{ arcsec pixels}$

$2 \text{ eV} @ 2\text{keV}$

$100\mu\text{s}$ decay time

Surrounding pixels:

$5 \times 5 \text{ arcsec pixels}$

$4 \text{ eV} @ 2 \text{ keV}$

$400\mu\text{s}$ decay time